

Solitons, Collapses and Turbulence



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SCT - 2017



Landau Institute for Theoretical Physics (Chernogolovka, Russia) , May 2017

*Как этот мир велик в лучах рабочей лампы,
А в памяти очах как бесконечно мал...*

(Ш.Бодлер, пер. М.Цветаевой)

100 years of General Relativity



Collision of gravitational and electromagnetic solitons with strong electromagnetic waves of arbitrary amplitudes and profiles in the expanding background space-time

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- ★ ***Gravitational and electromagnetic waves in (linear approximation)***
 - “Gertsenstein effect” (M.E. Gertsentein 1961)
 - mutual transformation of waves (GA & Sibgatullin 1974)
- ★ ***Nonlinear waves in the expanding universe (exact solutions):***
 - the symmetric Kasner space-time as a background for waves
 - travelling electromagnetic waves of arbitrary amplitudes
 - gravitational and electromagnetic solitons
 - collision of solitons with travelling waves

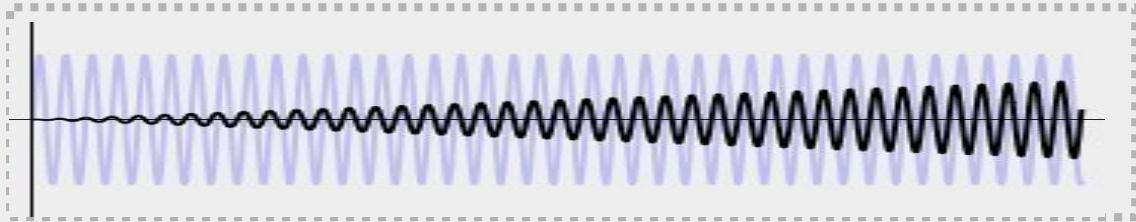
Einstein - Maxwell equations

$$\begin{cases} R_{ik} = -\frac{\varkappa}{4\pi} \left(F_{il} F_k{}^l - \frac{1}{4} g_{ik} F_{lm} F^{lm} \right) \\ \nabla_k F^{ik} = 0, \quad \nabla_{[i} F_{jk]} = 0 \end{cases}$$

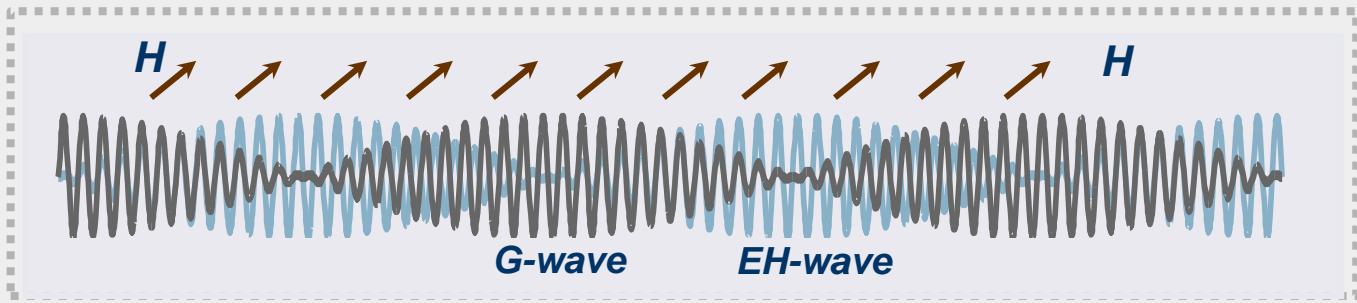
Gravitational and electromagnetic waves in external fields

(small amplitudes -- linear approximation):

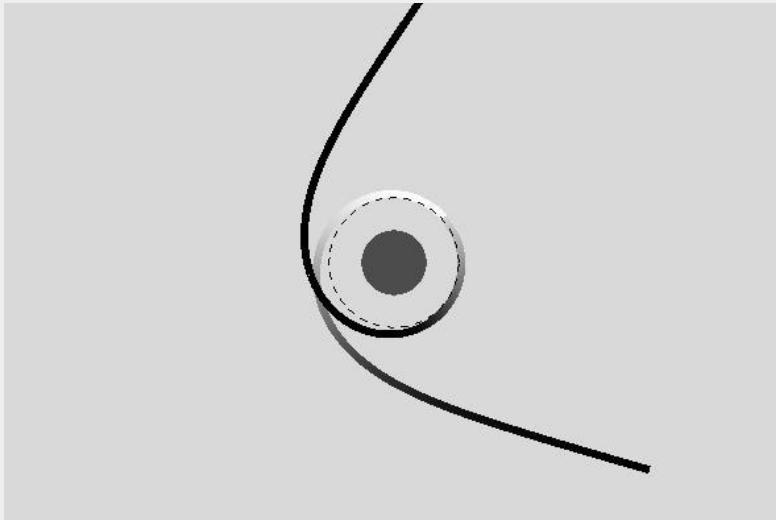
Gertsenstein M.E. (1961):



Sibgatullin N.R.
& GA (1974):



GA (1976):



Einstein - Maxwell equations

$$\begin{cases} R_{ik} = -\frac{\kappa}{4\pi} \left(F_{il} F_k{}^l - \frac{1}{4} g_{ik} F_{lm} F^{lm} \right) \\ \nabla_k F^{ik} = 0, \quad \nabla_{[i} F_{jk]} = 0 \end{cases}$$

Complex Ernst potentials

$$\begin{array}{ccc} g_{ik}(x^1, x^2) & \iff & \mathcal{E}(x^1, x^2), \\ F_{ik}(x^1, x^2) & & \Phi(x^1, x^2) \end{array}$$

Wave fields

$$u = t - x, \quad v = t + x$$

Complex Ernst equations:

$$\begin{cases} (\operatorname{Re} \mathcal{E} + \Phi \bar{\Phi}) \left(2\mathcal{E}_{uv} + \frac{\alpha_u}{\alpha} \mathcal{E}_v + \frac{\alpha_v}{\alpha} \mathcal{E}_u \right) - (\mathcal{E}_u + 2\bar{\Phi}\Phi_u) \mathcal{E}_v - (\mathcal{E}_v + 2\bar{\Phi}\Phi_v) \mathcal{E}_u = 0 \\ (\operatorname{Re} \mathcal{E} + \Phi \bar{\Phi}) \left(2\Phi_{uv} + \frac{\alpha_u}{\alpha} \Phi_v + \frac{\alpha_v}{\alpha} \Phi_u \right) - (\mathcal{E}_u + 2\bar{\Phi}\Phi_u) \Phi_v - (\mathcal{E}_v + 2\bar{\Phi}\Phi_v) \Phi_u = 0 \\ \alpha_{uv} = 0 \end{cases}$$

Inverse scattering approach to Einstein - Maxwell fields¹⁾

Dynamical equations $\mathbf{U}(\xi, \eta), \mathbf{V}(\xi, \eta)$

$$\left\{ \begin{array}{l} \partial_\eta \mathbf{U} + \partial_\xi \mathbf{V} + \frac{[\mathbf{U}, \mathbf{V}]}{\xi - \eta} = 0 \\ \partial_\eta \mathbf{U} - \partial_\xi \mathbf{V} = 0 \end{array} \right. \quad \parallel \quad \begin{array}{l} \mathbf{U} \cdot \mathbf{U} = i \mathbf{U}, \quad \text{Tr } \mathbf{U} = i \\ \mathbf{V} \cdot \mathbf{V} = i \mathbf{V}, \quad \text{Tr } \mathbf{V} = i \end{array}$$

Associated spectral problem $\Psi = \Psi(\xi, \eta, w)$

$$\left\{ \begin{array}{l} 2i(w - \xi) \partial_\xi \Psi = \mathbf{U}(\xi, \eta) \Psi \\ 2i(w - \eta) \partial_\eta \Psi = \mathbf{V}(\xi, \eta) \Psi \end{array} \right. \quad \parallel \quad \begin{array}{l} \mathbf{U} \cdot \mathbf{U} = i \mathbf{U} \quad \text{tr } \mathbf{U} = i \\ \mathbf{V} \cdot \mathbf{V} = i \mathbf{V} \quad \text{tr } \mathbf{V} = i \end{array}$$

Matrix integral $\mathbf{K}(w)$

$$\left\{ \begin{array}{l} \Psi^\dagger \mathbf{W} \Psi = \mathbf{K}(w) \\ \mathbf{K}^\dagger(w) = \mathbf{K}(w) \end{array} \right. \quad \parallel \quad \frac{\partial \mathbf{W}}{\partial w} = \Omega, \quad \Omega = \begin{pmatrix} 0 & 1 & 0 \\ -1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

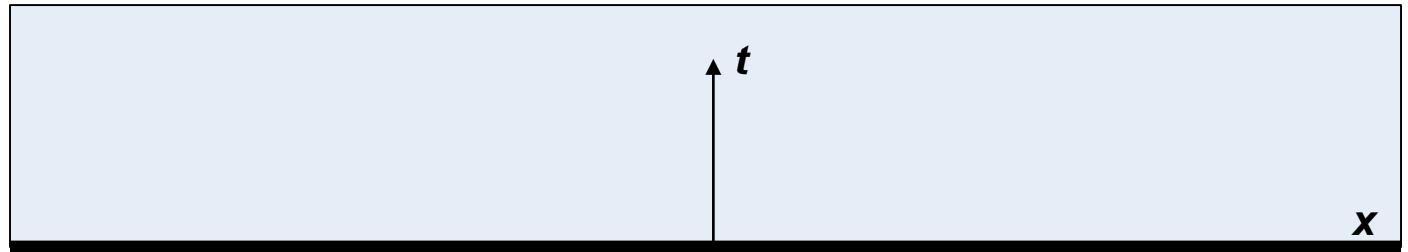
“Dressing” method for constructing of solitons: $(\overset{o}{g}_{ab}, \overset{o}{A}_b, \overset{o}{\Psi}) \rightarrow (g_{ab}, A_b \Psi)$

$$\Psi = \chi \cdot \overset{o}{\Psi}, \quad \chi = \mathbf{I} + \sum_{k=1}^N \frac{\mathbf{R}_k(\xi, \eta)}{w - w_k}, \quad \chi^{-1} = \mathbf{I} + \sum_{k=1}^N \frac{\mathbf{S}_k(\xi, \eta)}{w - \tilde{w}_k}$$

Background for waves: the symmetric Kazner vacuum universe

$$ds^2 = -\frac{1}{\sqrt{t}}(dt^2 - dx^2) + t(dy^2 + dz^2) \quad \xi = -x + t = u \\ \eta = -x - t = -v$$

Space-time:



Solution of the spectral problem:

$$\Psi = \frac{1}{2} \left(\frac{w - \xi_0}{w - \xi} \right)^{1/2} \begin{pmatrix} 1 \\ it_0 \end{pmatrix} \otimes \begin{pmatrix} 1, -\frac{i}{t_0} \end{pmatrix} + \frac{1}{2} \left(\frac{w - \eta_0}{w - \eta} \right)^{1/2} \begin{pmatrix} 1 \\ -it_0 \end{pmatrix} \otimes \begin{pmatrix} 1, \frac{i}{t_0} \end{pmatrix}$$

$$\Psi^{-1} = \frac{1}{2} \left(\frac{w - \xi}{w - \xi_0} \right)^{1/2} \begin{pmatrix} 1 \\ it_0 \end{pmatrix} \otimes \begin{pmatrix} 1, -\frac{i}{t_0} \end{pmatrix} + \frac{1}{2} \left(\frac{w - \eta}{w - \eta_0} \right)^{1/2} \begin{pmatrix} 1 \\ -it_0 \end{pmatrix} \otimes \begin{pmatrix} 1, \frac{i}{t_0} \end{pmatrix}$$

Gravitational and electromagnetic soliton waves on the symmetric Kasner background

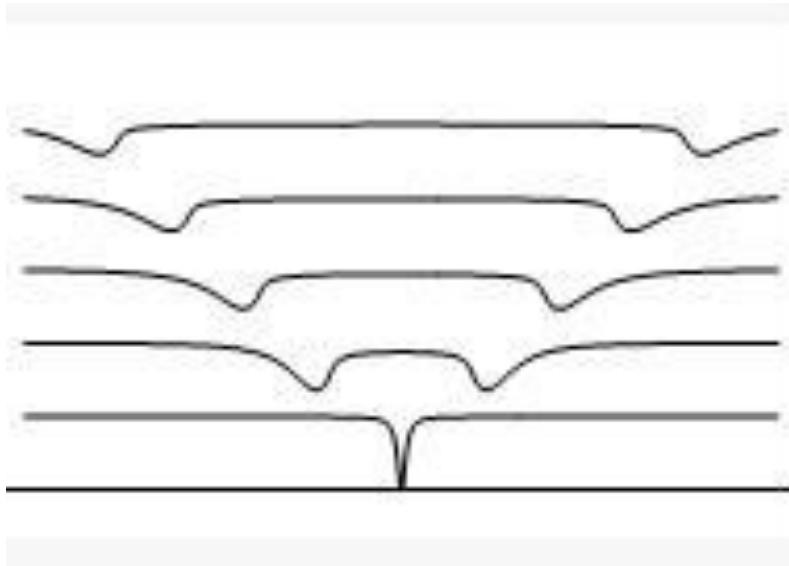
$$\left\{ \begin{array}{l} ds^2 = -f du dv + t \left[h_{yy} dy^2 + 2h_{yz} dy dz + h_{zz} dz^2 \right], \\ A_i = \{0, 0, A_y, A_z\}, \end{array} \right. \quad \left| \begin{array}{l} \ell + is, \\ c, \\ d \end{array} \right.$$

$$\begin{aligned} h_{yy} &= \frac{(\mathcal{G} - 4isc)(\bar{\mathcal{G}} + 4is\bar{c})}{\mathcal{D}\bar{\mathcal{D}}}, & h_{yz} &= -\frac{4s(c\bar{\mathcal{G}} + \bar{c}\mathcal{G})}{\mathcal{D}\bar{\mathcal{D}}}, & \{A_y, A_z\} &= \Re\{\Phi, \widetilde{\Phi}\}, \\ h_{zz} &= \frac{(\mathcal{G} + 4isc)(\bar{\mathcal{G}} - 4is\bar{c})}{\mathcal{D}\bar{\mathcal{D}}}, & f &= \frac{1}{\sqrt{t}} \cdot \frac{\mathcal{D}\bar{\mathcal{D}}}{4r_+ r_-} & \Phi &= -2sd\mathcal{F}/\mathcal{D}, \\ & & & & \widetilde{\Phi} &= 2isd\tilde{\mathcal{F}}/\mathcal{D} \end{aligned}$$

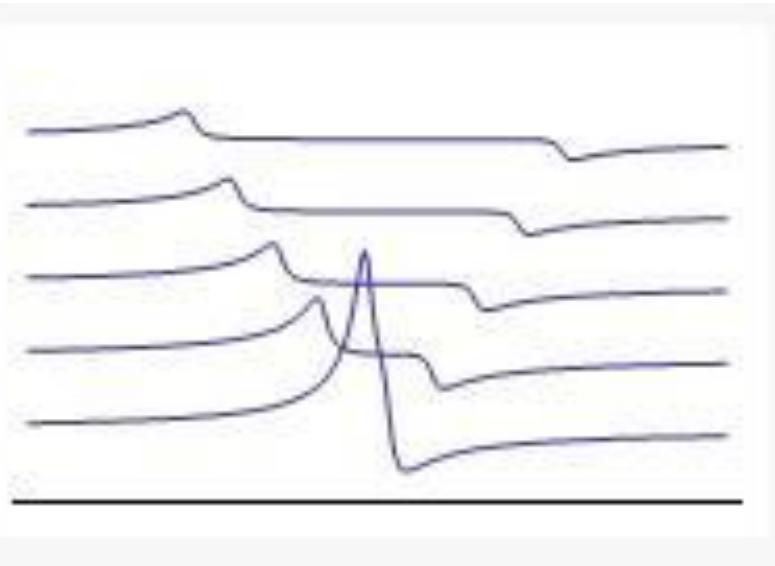
$$\begin{aligned} \mathcal{D} &= \bar{K}_+ K_- - c\bar{c}K_+\bar{K}_- + d\bar{d}\bar{K}_+\bar{K}_-, & K_+ &= \sqrt{R_+} + \frac{is}{\sqrt{R_+}}, \\ \mathcal{G} &= K_+ K_- - c\bar{c}\bar{K}_+\bar{K}_- + d\bar{d}K_+\bar{K}_-, & K_- &= \sqrt{R_-} + \frac{is}{\sqrt{R_-}}, \\ \mathcal{F} &= \bar{K}_+ - \bar{c}\bar{K}_-, & \tilde{\mathcal{F}} &= \bar{K}_+ + \bar{c}\bar{K}_-. \end{aligned}$$

$$\begin{aligned} R_+ &= \ell - u + r_+, & r_+ &= \sqrt{(\ell - u)^2 + s^2}, \\ R_- &= \ell + v + r_-, & r_- &= \sqrt{(\ell + v)^2 + s^2}. \end{aligned}$$

One-soliton solution on the symmetric Kazner background



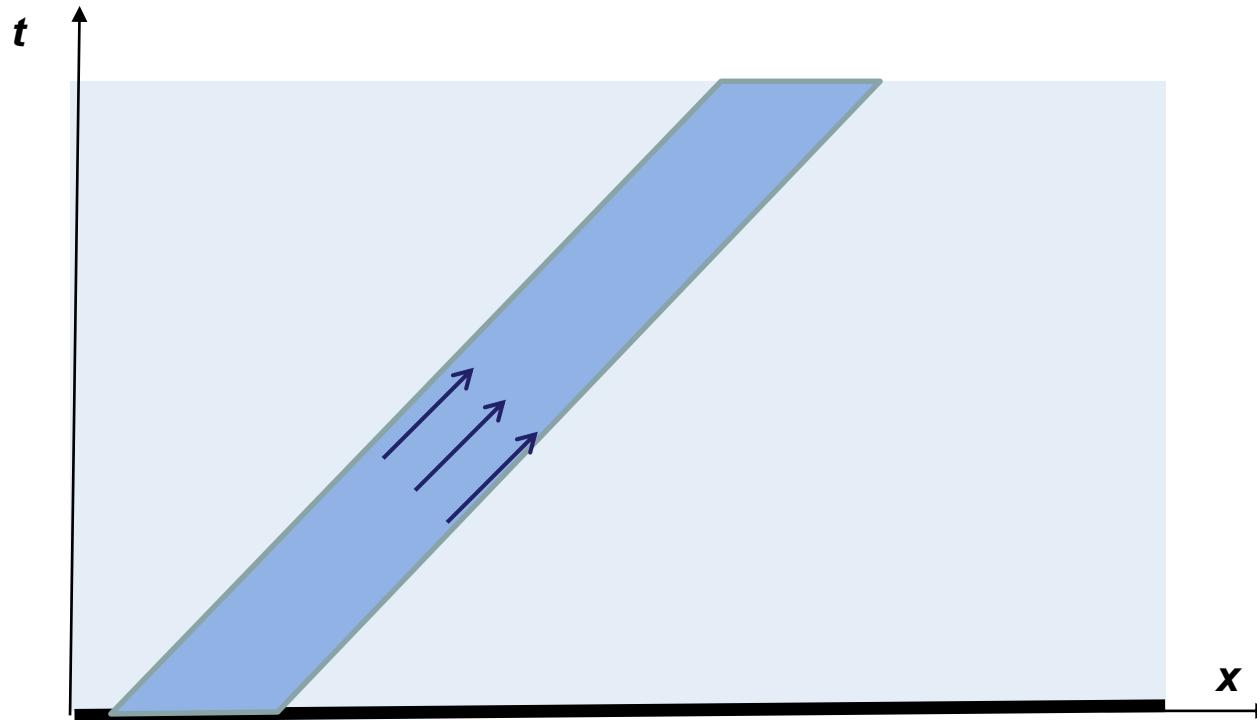
$\text{Im } c = 0, \ d = 0$



$\text{Re } c = 0, \ d = 0$

Plane electromagnetic waves on the symmetric Kazner background¹⁾

$$\left\{ \begin{array}{l} ds^2 = -\frac{1}{\sqrt{t}} e^{4\mathcal{B}(u)} du dv + t(dy^2 + dz^2), \\ A_i = \{0, 0, \phi(u), 0\}, \quad \mathcal{B}(u) = \int \phi'^2(u) du \end{array} \right.$$



Solutions of the linear system for a travelling electromagnetic wave

$$\Psi_L = \begin{pmatrix} 1 & e^{S_+(\xi, w)} & e^{-S_+(\xi, w)} \\ \frac{i}{\sqrt{w-\eta}} & \frac{\sqrt{w-\xi}}{ie^{S_+(\xi, w)}} & \frac{\sqrt{w-\xi}}{ie^{-S_+(\xi, w)}} \\ -\frac{i}{\sqrt{w-\eta}} & \frac{ie^{S_+(\xi, w)}}{\sqrt{w-\xi}} & \frac{ie^{-S_+(\xi, w)}}{\sqrt{w-\xi}} \\ 0 & 2i\sqrt{2}e^{S_+(\xi, w)} \left(1 - \frac{\phi_+(\xi)}{\sqrt{w-\xi}}\right) & -2i\sqrt{2}e^{-S_+(\xi, w)} \left(1 + \frac{\phi_+(\xi)}{\sqrt{w-\xi}}\right) \end{pmatrix}$$

$$S_+(\xi, w) = \int_{\xi_0}^{\xi} \frac{\phi'_+(\chi)}{\sqrt{w-\chi}} d\chi$$

Nonlinear interaction of gravitational and electromagnetic solitons with a non-soliton electromagnetic wave of arbitrary profile ¹⁾

Solution parameters: $\ell + is, c, d$ and $\phi(u)$

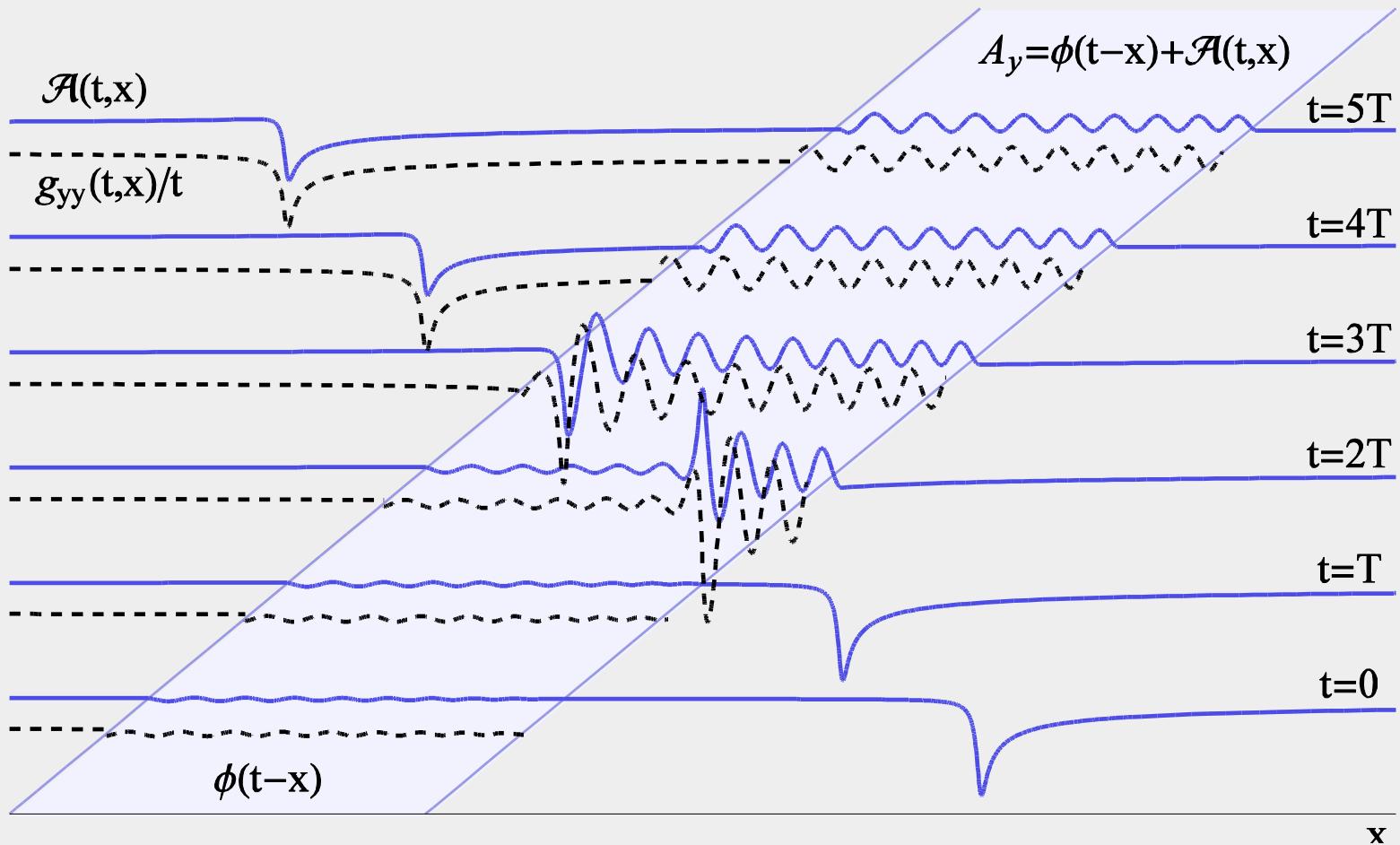
**Behaviour of a soliton in
the interaction region:** $c \rightarrow \hat{c}(u), \quad \Phi \rightarrow \Phi + \phi(u),$
 $d \rightarrow \hat{d}(u), \quad \tilde{\Phi} \rightarrow \tilde{\Phi} + i\phi(u), \quad f \rightarrow e^{4\mathcal{B}(u)} f$

Evolution of the soliton parameters:

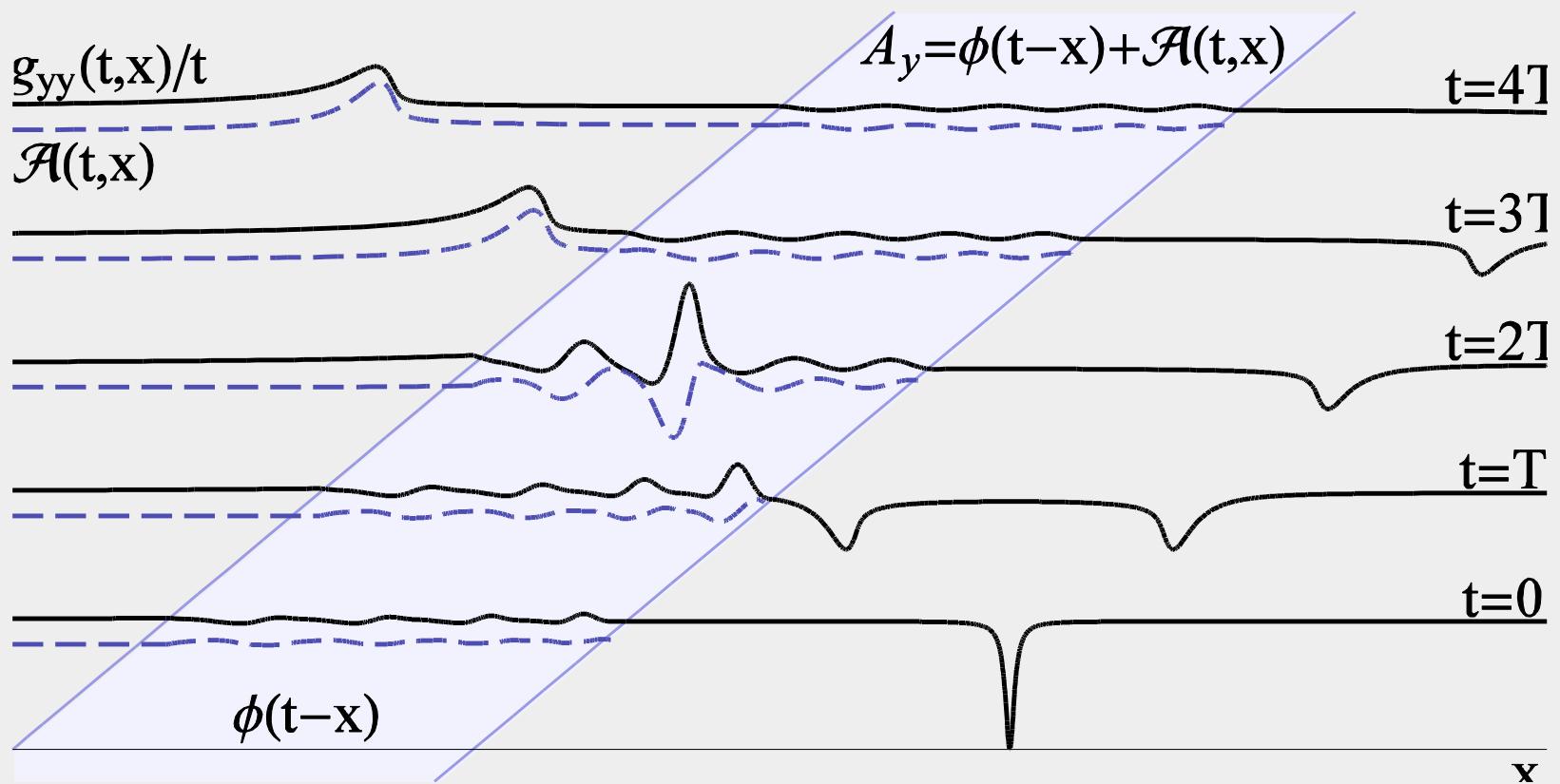
$$\begin{aligned}\hat{c} &= c \cosh S(u) - id \sinh S(u), \\ \hat{d} &= ic \sinh S(u) + d \cosh S(u),\end{aligned}$$

$$S(u) = \sqrt{2} \int_{u_1}^u \frac{\phi'(u) du}{\sqrt{w_o - u}}$$

Collision of electromagnetic soliton with non-soliton electromagnetic wave



Collision of vacuum soliton with travelling electromagnetic wave



An aerial photograph of a vast expanse of white, fluffy cumulus clouds against a clear blue sky. The clouds are more dense in the foreground and appear to stretch to the horizon.

Thank you